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## FADDEEV GRAVITY ACTION ON THE PIECEWISE CONSTANT FUNDAMENTAL VECTOR FIELDS

V. M. Khatsymovsky

Budker Institute of Nuclear Physics of Siberian Branch Russian Academy of Sciences, Novosibirsk, 630090, Russia.

E-mail: khatsym@gmail.com

In the Faddeev formulation of gravity, the metric is regarded as composite field, bilinear of  $d = 10$  4-vector fields. We derive the minisuperspace (discrete) Faddeev action by evaluating the Faddeev action on the spacetime composed of the (flat) 4-simplices with constant 4-vector fields. This is an analog of the Regge action obtained by evaluating the Hilbert-Einstein action on the spacetime composed of the flat 4-simplices. One of the new features of this formulation is that the simplices are not required to coincide on their common faces. Also an analog of the Barbero-Immirzi parameter  $\gamma$  can be introduced in this formalism.

**Keywords:** Einstein gravity, Regge calculus, composite metric, Faddeev gravity, discrete gravity.

### 1 Introduction

Regge calculus [1] is a minisuperspace formulation of gravity, that is, exact general relativity (GR) on a family of the metric fields (piecewise flat ones) sufficiently large to approximate any metric with any accuracy. Spacetime is taken as the set of the flat 4d tetrahedra. Let  $\sigma^n$  be  $n$ -dimensional tetrahedron or  $n$ -simplex. Then the Einstein action is a sum over triangles  $\sigma^2$ ,

$$\frac{1}{2} \int R \sqrt{g} d^4x = \sum_{\sigma^2} A_{\sigma^2} \alpha_{\sigma^2}. \quad (1)$$

Here,  $A_{\sigma^2}$  is area of the triangle  $\sigma^2$ ,  $\alpha_{\sigma^2}$  is the defect angle on the triangle  $\sigma^2$ . Regge calculus is expected to be the base for quantizing GR [2–4] difficult to implement by the other methods because of the formal nonrenormalizability of GR.

Let us consider some another minisuperspace theory based on the Faddeev formulation of gravity [5]. In the Faddeev gravity, the metric is composed of  $d = 10$  4-vector fields,

$$g_{\lambda\mu} = f_{\lambda}^A f_{\mu A}. \quad (2)$$

Here,  $\lambda, \mu, \dots = 1, 2, 3, 4$ ;  $A, B, \dots = 1, \dots, d$ . Simple example: locally, 4d Riemannian space can be considered as a hyper-surface in the 10d Euclidean space. If  $f^A(x)$  were its coordinates, then we would have

$$f_{\lambda}^A = \partial f^A / \partial x^{\lambda}. \quad (3)$$

In the Faddeev gravity, however,  $f_{\lambda}^A(x)$  are independent variables. We can introduce an alternative connection

$$\tilde{\Gamma}_{\lambda\mu\nu} = f_{\lambda}^A f_{\mu A, \nu} (f_{\mu A, \nu} \equiv \partial_{\nu} f_{\mu A}), \quad \tilde{\Gamma}_{\mu\nu}^{\lambda} = g^{\lambda\rho} \tilde{\Gamma}_{\rho\mu\nu}, \quad (4)$$

its curvature is

$$\begin{aligned} K_{\mu\nu\rho}^{\lambda} &= \tilde{\Gamma}_{\mu\rho, \nu}^{\lambda} - \tilde{\Gamma}_{\mu\nu, \rho}^{\lambda} + \tilde{\Gamma}_{\sigma\nu}^{\lambda} \tilde{\Gamma}_{\mu\rho}^{\sigma} - \tilde{\Gamma}_{\sigma\rho}^{\lambda} \tilde{\Gamma}_{\mu\nu}^{\sigma} \\ &= \Pi^{AB} (f_{A, \nu}^{\lambda} f_{\mu B, \rho} - f_{A, \rho}^{\lambda} f_{\mu B, \nu}). \end{aligned} \quad (5)$$

Here,

$$\Pi_{AB} = \delta_{AB} - f_A^{\lambda} f_{\lambda B} \quad (6)$$

is a projector. Note that it makes usual and covariant derivatives equivalent,

$$\Pi^{AB} f_{\lambda B, \mu} = \Pi^{AB} \nabla_{\mu} f_{\lambda B}. \quad (7)$$

Then we can write out the Faddeev action,

$$\begin{aligned} S &= \int g^{\lambda\nu} g^{\mu\rho} K_{\lambda\mu\nu\rho} \sqrt{g} d^4x + \frac{1}{2\gamma} \int \epsilon^{\lambda\mu\nu\rho} K_{\lambda\mu\nu\rho} d^4x \\ &= \int \Pi^{AB} \left[ (f_{A, \lambda}^{\lambda} f_{B, \mu}^{\mu} - f_{A, \mu}^{\lambda} f_{B, \lambda}^{\mu}) \sqrt{g} \right. \\ &\quad \left. - \frac{1}{\gamma} \epsilon^{\lambda\mu\nu\rho} f_{\lambda A, \mu} f_{\nu B, \rho} \right] d^4x. \end{aligned} \quad (8)$$

Here we have generalized it by adding the P-odd  $1/\gamma$ -term where  $\gamma$  is an analog [6] of the Barbero-Immirzi parameter [7–10] in the usual Einstein gravity (in the connection representation).

The variation of the action reads

$$\begin{aligned} \frac{\delta S}{2\sqrt{g} \delta f_A^{\lambda}} &= f_{\mu}^A \left( K_{\lambda}^{\mu} - \frac{1}{2} \delta_{\lambda}^{\mu} K - \frac{\epsilon^{\mu\nu\rho\sigma}}{2\gamma\sqrt{g}} K_{\lambda\nu\rho\sigma} \right) \\ &+ \Pi^{AB} \left[ f_{B, \nu}^{\nu} T_{\lambda\mu}^{\mu} + f_{B, \mu}^{\nu} T_{\nu\lambda}^{\mu} + f_{B, \lambda}^{\nu} T_{\mu\nu}^{\mu} \right. \\ &\quad \left. + \frac{\epsilon^{\tau\mu\nu\rho}}{2\gamma\sqrt{g}} (g_{\lambda\sigma} g_{\kappa\tau} - g_{\lambda\tau} g_{\kappa\sigma}) f_{B, \rho}^{\kappa} T_{\mu\nu}^{\sigma} \right]. \end{aligned} \quad (9)$$

Here, the  $\Pi^{AB}$ -part of the field eqs  $\delta S = 0$  turns out to be a linear homogeneous system for torsion  $T_{\mu\nu}^{\lambda} = f_A^{\lambda} (f_{\mu, \nu}^A - f_{\nu, \mu}^A)$  with a nondegenerate  $24 \times 24$

matrix leading to  $T_{\mu\nu}^\lambda = 0$ . This gives  $\tilde{\Gamma}_{\mu\nu}^\lambda = \Gamma_{\mu\nu}^\lambda$ , the Cristoffel. Then  $K_{\mu\nu\rho}^\lambda = R_{\mu\nu\rho}^\lambda$ , the Riemann for which  $\epsilon^{\mu\nu\rho\sigma} R_{\lambda\nu\rho\sigma} \equiv 0$ , and the  $f_\mu^A$ -part of the field eqs turns out to be the Einstein eqs.

### 2 The piecewise constant fields

In GR, taking the metric as

$$ds^2 = g_{nn}(dx^n)^2 + g_{\alpha\beta}dx^\alpha dx^\beta, \quad (10)$$

where  $n$  is any of 1, 2, 3, 4, and  $\alpha, \beta, \gamma, \dots \neq n$ , we can separate out the derivatives over  $x^n$  squared in the action as follows,

$$\int R\sqrt{g}d^4x = \frac{1}{4} \int (g^{\alpha\gamma}g^{\beta\delta} - g^{\alpha\beta}g^{\gamma\delta})g_{\alpha\beta,n}g_{\gamma\delta,n}g^{nn}\sqrt{g}d^4x + \dots \quad (11)$$

If the metric is taken to be piecewise constant (with  $g_{\lambda\mu} = const$  inside the 4-simplices), this implies continuity condition for the induced on the faces metric,

$$g_{\alpha\beta}|_{x^n=x_0^n+0} - g_{\alpha\beta}|_{x^n=x_0^n-0} = 0 \quad \forall x_0^n, \quad (12)$$

if some 3-face is described locally by  $x^n = x_0^n$ . That is, the simplices should coincide on their common faces.

In the Faddeev gravity, the fields  $f_\lambda^A$  can be taken to be piecewise constant. The action

$$S = \int \Pi^{AB} \left[ (f_{A,\lambda}^\lambda f_{B,\mu}^\mu - f_{A,\mu}^\lambda f_{B,\lambda}^\mu) \sqrt{g} - \frac{1}{\gamma} \epsilon^{\lambda\mu\nu\rho} f_{\lambda A,\mu} f_{\nu B,\rho} \right] d^4x \quad (13)$$

does not contain the square of any derivative and does not require continuity condition on  $f_\lambda^A(x)$  and thus on the metric  $g_{\lambda\mu} = f_\lambda^A f_{\mu A}$ . The simplices are not required to coincide on their common faces, and we can take  $f_\lambda^A(x) = const$  independently in these 4-simplices.

### 3 The discrete Faddeev action

To evaluate the discrete action [11], we divide  $\mathbb{R}^4$  as the set of points  $x = (x^1, x^2, x^3, x^4)$  by the hypersurfaces  $a_\lambda x^\lambda + b = 0$  (mathematical hyperplanes) into polytopes, in particular, 4-simplices.  $f_\lambda^A$  are independent in these 4-simplices.

Then we note that the expression

$$Q_{AB} = f_{A,\lambda}^\lambda f_{B,\mu}^\mu - f_{A,\lambda}^\mu f_{B,\mu}^\lambda \quad (14)$$

1) is zero if  $f_\lambda^A$  depends only on one coordinate: contribution of the 3-simplices vanishes, and  $Q_{AB} = const \cdot \delta(x^1)\delta(x^2)$  in a neighborhood of the 2-simplex

$\sigma^2$  at  $x^1 = 0, x^2 = 0$  (triangle in  $x^3, x^4$ -plane); 2) is a full derivative

$$Q_{AB} = \partial_\lambda (f_A^\lambda \partial_\mu f_B^\mu - f_A^\mu \partial_\mu f_B^\lambda) \quad (15)$$

so that

$$\int Q_{AB} dx^1 dx^2 = \oint_C (f_A^1 df_B^2 - f_A^2 df_B^1), \quad (16)$$

where the contour  $C$  encircles the triangle  $\sigma^2$ . This gives

$$2Q_{AB} = \delta(x^1)\delta(x^2) \sum_{i=1}^n [f_A^1(\sigma_i^4) f_B^2(\sigma_{i+1}^4) - f_A^1(\sigma_{i+1}^4) f_B^2(\sigma_i^4)] + (A \leftrightarrow B), \quad (17)$$

where the 4-simplices  $\sigma_i^4, i = 1 \div n$  surround the triangle  $\sigma^2$ . The action is

$$\frac{1}{2} \Pi^{AB}(\sigma^2) \sum_{i=1}^n \left\{ [f_A^1(\sigma_i^4) f_B^2(\sigma_{i+1}^4) - f_A^1(\sigma_{i+1}^4) f_B^2(\sigma_i^4)] \sqrt{g(\sigma^2)} - \frac{1}{\gamma} [f_{4A}(\sigma_i^4) f_{3B}(\sigma_{i+1}^4) - f_{4A}(\sigma_{i+1}^4) f_{3B}(\sigma_i^4)] \right\} \cdot (\Delta x_{\sigma_3^4}^3 \Delta x_{\sigma_4^4}^4 - \Delta x_{\sigma_4^4}^3 \Delta x_{\sigma_3^4}^4), \quad (18)$$

where  $\sigma^2$  is formed by the edges  $\sigma_3^1, \sigma_4^1$  with the 4-vectors  $\Delta x_{\sigma_3^4}^\lambda, \Delta x_{\sigma_4^4}^\lambda$ .

### 4 An invariant form of the action

We can rewrite this action in some invariant variables. This invariance is that one w. r. t. the choice of the coordinates  $x^\lambda$  assigned to the vertices. Choose some two edges  $\sigma_1^1, \sigma_2^1$  in addition to the above  $\sigma_3^1, \sigma_4^1$  spanning  $\sigma^2$  so that  $\sigma_1^1, \sigma_2^1, \sigma_3^1, \sigma_4^1$  span some  $\sigma_{i_0}^4 \supset \sigma^2$ . The difference between  $x^\lambda$  of the endpoints of  $\sigma_\mu^1$  is  $\Delta x_{\sigma_\mu^1}^\lambda$ . Invariant edge variables  $f_A^{\sigma_\mu^1}$  or  $f_{\sigma_\mu^1}^A$  are defined by

$$f_A^\lambda = \sum_\mu f_A^{\sigma_\mu^1} \Delta x_{\sigma_\mu^1}^\lambda, \quad f_{\sigma_\lambda^1}^A = f_\mu^A \Delta x_{\sigma_\lambda^1}^\mu. \quad (19)$$

Then we can write the contribution of  $\sigma^2$  to the action,

$$\frac{1}{2} \Pi^{AB}(\sigma^2) \sum_{i=1}^n \left\{ [f_A^{\sigma_1^1}(\sigma_i^4) f_B^{\sigma_2^1}(\sigma_{i+1}^4) - f_A^{\sigma_1^1}(\sigma_{i+1}^4) f_B^{\sigma_2^1}(\sigma_i^4)] \sqrt{\det \|g_{\sigma_\lambda^1 \sigma_\mu^1}\|} - \frac{1}{\gamma} [f_{\sigma_4^1 A}(\sigma_i^4) f_{\sigma_3^1 B}(\sigma_{i+1}^4) - f_{\sigma_4^1 A}(\sigma_{i+1}^4) f_{\sigma_3^1 B}(\sigma_i^4)] \right\}, \quad (20)$$

where  $\Pi_B^A = \delta_B^A - \sum_\lambda f_{\sigma_\lambda^1}^A f_{\sigma_\lambda^1}^B, g_{\sigma_\lambda^1 \sigma_\mu^1} = f_{\sigma_\lambda^1}^A f_{\sigma_\mu^1 A}$  (metric edge components).

## 5 Restoring the continuum action from the discrete minisuperspace one

We can obtain the continuum Faddeev action from the found discrete minisuperspace one. This reverse transition is intended to show that the information encoded in the minisuperspace formulation is sufficient to reproduce the essential degrees of freedom of the continuum theory. This procedure is analogous to the work [12] for the usual GR where the Regge action has been shown to tend to the Hilbert-Einstein one if simplicial decomposition of the given smooth manifold is made finer and finer. To do this, we choose the fields to approximate some fixed smooth  $f_\lambda^A(x)$  arbitrarily closely by making decomposition of spacetime into the 4-simplices finer and finer (the coordinate steps along the edges  $\Delta x_{\sigma_1}^\lambda \rightarrow 0$ ). Then we choose the discrete variable  $f_A^\lambda(\sigma^4)$  to be  $f_A^\lambda(x)$  at  $x = x_{\sigma^4}$ , some central point in  $\sigma^4$ . The procedure is most easily illustrated by the case  $n = 4$ :

$$\begin{aligned} \sum_{i=1}^4 [f_A^1(\sigma_i^4) f_B^2(\sigma_{i+1}^4) - f_A^1(\sigma_{i+1}^4) f_B^2(\sigma_i^4)] \\ = \Delta_u f_A^1 \Delta_v f_B^2 - \Delta_v f_A^1 \Delta_u f_B^2 \end{aligned} \quad (21)$$

where  $\Delta_u f \equiv f(\sigma_3^4) - f(\sigma_1^4)$  and  $\Delta_v f \equiv f(\sigma_4^4) - f(\sigma_2^4)$  in the continuum limit become derivatives. This gives us the basic structure (some two-dimensional Jacobian) which being summed over the different contravariant components  $f_{\dots}^\lambda$  and variations between the different pairs of simplices  $\Delta f = f(\sigma^4) - f(\sigma^{4'})$  just leads to  $f_{A,\lambda}^\lambda f_{B,\mu}^\mu - f_{A,\lambda}^\mu f_{B,\mu}^\lambda$  and thus to the Faddeev action. The  $1/\gamma$ -part of the action is considered completely analogously.

## 6 Possibility of the cubic decomposition of spacetime

A new interesting property of the Faddeev formulation as compared to the Regge calculus is the possibility to decompose the curved spacetime into the (flat) cubes. Consider decomposition of  $\mathbb{R}^4$  into the rectangular parallelepipeds or cuboids

$$\begin{aligned} \{x|x_0^\lambda + n^\lambda \varepsilon^\lambda \leq x^\lambda \\ \leq x_0^\lambda + (n^\lambda + 1)\varepsilon^\lambda, \lambda = 1, 2, 3, 4\}. \end{aligned} \quad (22)$$

Then  $f_A^\lambda(x)$  is assumed to be constant in each of these cuboids. Above derivation of the minisuperspace Faddeev action remains the same if  $\sigma^4$  are not 4-simplices but cuboids. In Regge calculus, the continuity of metric on the faces leads to that the spacetime composed of flat cuboids is flat as well. In the Faddeev gravity, continuity of  $f_A^\lambda(x)$  is not required, and the flat cuboids can be used for modeling the curved spacetime. The cuboid action looks just as the naively discretized action or the continuum one (8) in which the derivatives are substituted by their finite difference counterparts. It is much simpler than the simplicial one and, at the same time, it is a minisuperspace action.

## 7 Conclusion

To summarize, we have constructed some discrete formalism analogous to the Regge calculus, but based not on the usual GR, but on its Faddeev form. Some new features of this formalism are the following.

- 1) The invariant physical variables are edge 10-vectors  $f_{\sigma^1}^A(\sigma^4)$  (or  $f_{\sigma^1}^A(\sigma^4)$ ), independent for the different 4-simplices  $\sigma^4$  containing the edge  $\sigma^1$ .
- 2) The 4-simplices do not necessarily coincide on their common faces.
- 3)  $1/\gamma$ -term is present which is an analogue of the so called Barbero-Immirzi parity odd term in the connection representation of the GR action.

An advantage of the discrete Faddeev formalism is the possibility to use cubic decomposition instead of the simplicial one for modeling the curved spacetime due to possibility of the field  $f_A^\lambda$  discontinuities in the Faddeev gravity. The cuboid action looks simple as the naive discretization of the continuum one.

Also this possibility of the field  $f_A^\lambda$  discontinuities allows one to evaluate the area spectrum of any surface as the sum of spectra of independent triangles. In blackhole physics, reasonable physical arguments require a discrete area spectrum, but up to now, only one gravity theory (LQG) allows to obtain it [13–15]. We hope that the discrete Faddeev gravity approach is also able to handle this task.

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*В. М. Хацимовский*

## ГРАВИТАЦИОННОЕ ДЕЙСТВИЕ ФАДДЕЕВА НА КУСОЧНО-ПОСТОЯННЫХ ФУНДАМЕНТАЛЬНЫХ ВЕКТОРНЫХ ПОЛЯХ

В формулировке Фаддеева гравитации метрика считается композитным полем, билинейным по  $d = 10$  4-векторным полям. Мы выводим (дискретное) действие Фаддеева в конфигурационном минисуперпространстве путем вычисления действия Фаддеева на пространстве-времени, составленном из (плоских) 4-симплексов с постоянными 4-векторными полями. Это аналог действия Редже, полученного вычислением действия Гильберта-Эйнштейна на пространстве-времени, составленном из плоских 4-симплексов. Одна из новых черт этой формулировки состоит в том, что симплексы не обязаны совпадать на общих гранях. Также можно ввести в этом формализме аналог параметра Барbero-Иммирци  $\gamma$ .

**Ключевые слова:** эйнштейновская гравитация, исчисление Редже, композитная метрика, гравитация Фаддеева, дискретная гравитация.

Хацимовский В. М., доктор физико-математических наук.  
**Институт ядерной физики им. Будкера СО РАН.**  
Пр. Лаврентьева, 11, 630090 Новосибирск, Россия.  
E-mail: khatsym@gmail.com